



Pygmy Dipole Resonance (PDR) in exotic ^{68}Ni



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Introduction

Pygmy Resonance

Analysis and Results

RISING data on ^{68}Ni

Virtual photon scattering

Conclusions



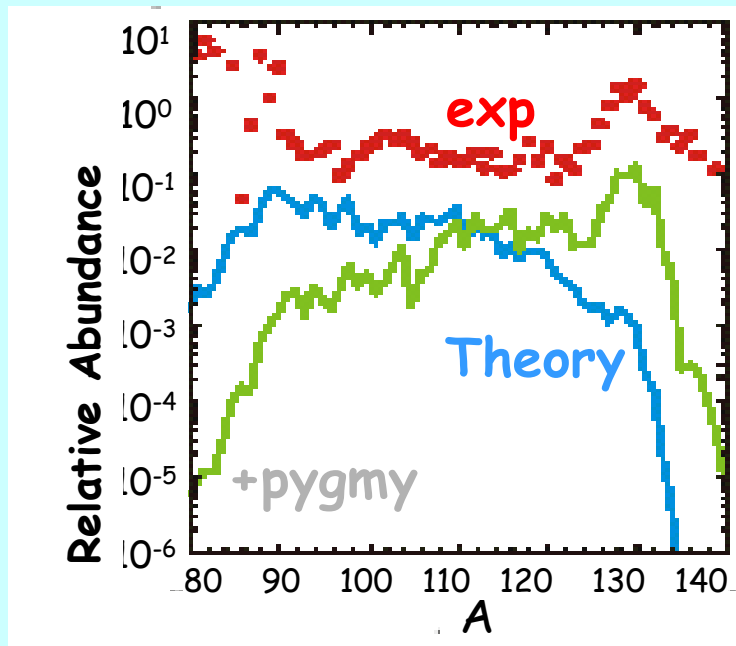
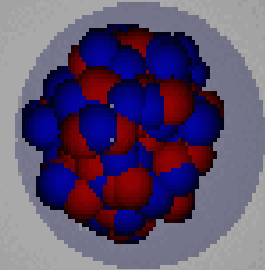
EGP-Workshop European Gammapool Workshop - Paris, May 27th-30th 2008



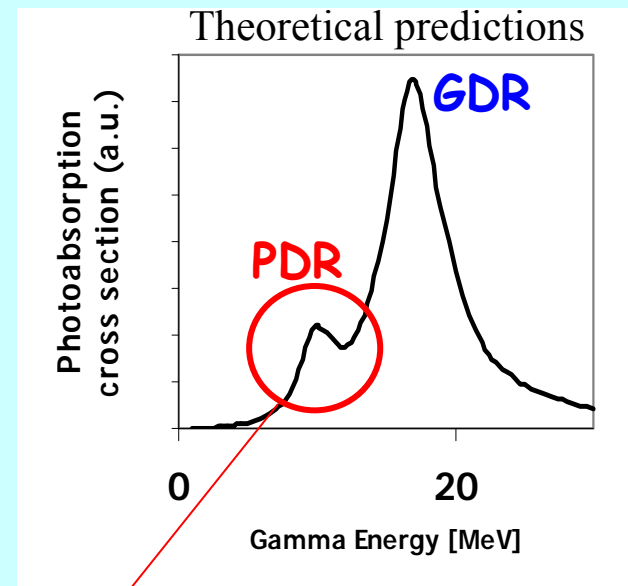
Pygmy Dipole Resonance in n-rich nuclei

- ▶ how collective properties change with n number
- ▶ important astrophysical implications for r-process

n excess vs inert core



S. Goriely, Phys. Lett. B 436 (1998) 10



E1 strength **shifted** towards low energy

collective or non-collective nature of E1 states ?

Virtual photon scattering technique

First pygmy coulomb excitation experiment with a fast relativistic beam

GDR - PYGMY Excitation

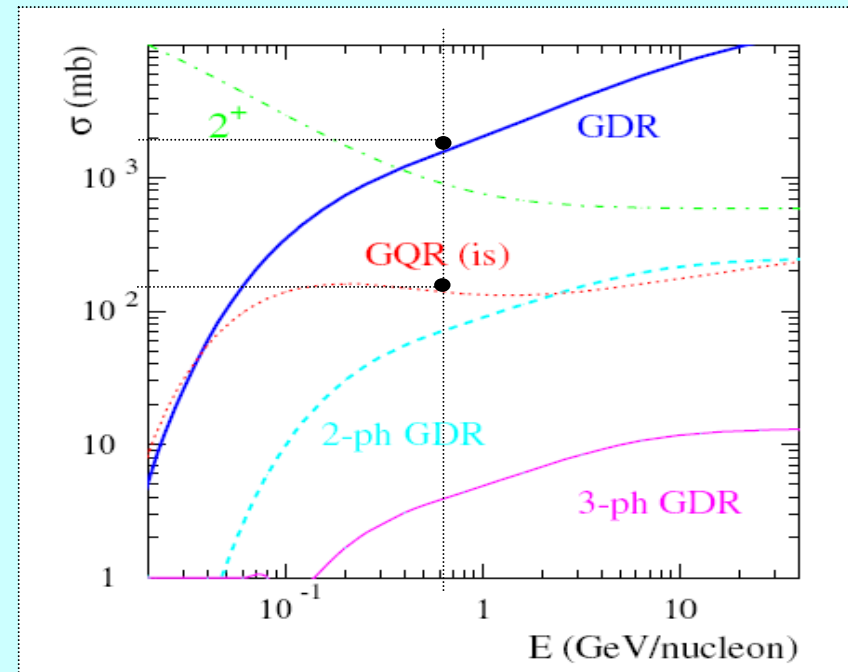
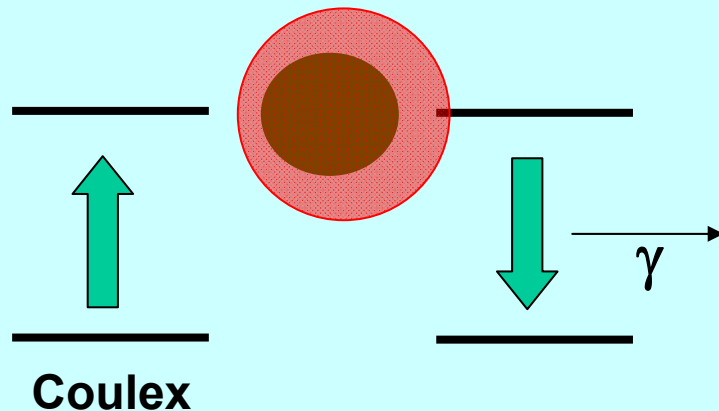
600 MeV/u ^{68}Ni + ^{197}Au (April 2005)

400 MeV/u ^{68}Ni + ^{197}Au (May 2004)

- High selectivity for dipole excitation

$$\frac{\sigma(GDR)}{\sigma(GQR)} \approx 20$$

Virtual photon excitation
and decay of **GDR - PYGMY**



T.Aumann et al EPJ 26(2005)441

→ At large energies the cross section for the Coulomb excitation of the GR overcomes the nuclear geometrical cross section!

**GDR Ground state decay
branching ratio
~ 2% measured on ^{208}Pb**

[Beene et al PRC 41(1990)920]

Coulomb excitation of ^{68}Ni @ 600 AMeV

RISING ARRAY

Euroball **15 Clusters**

Located at 16.5° , 33° , 36° degrees
Energetic threshold ~ 100 keV

Hector **8 BaF₂**

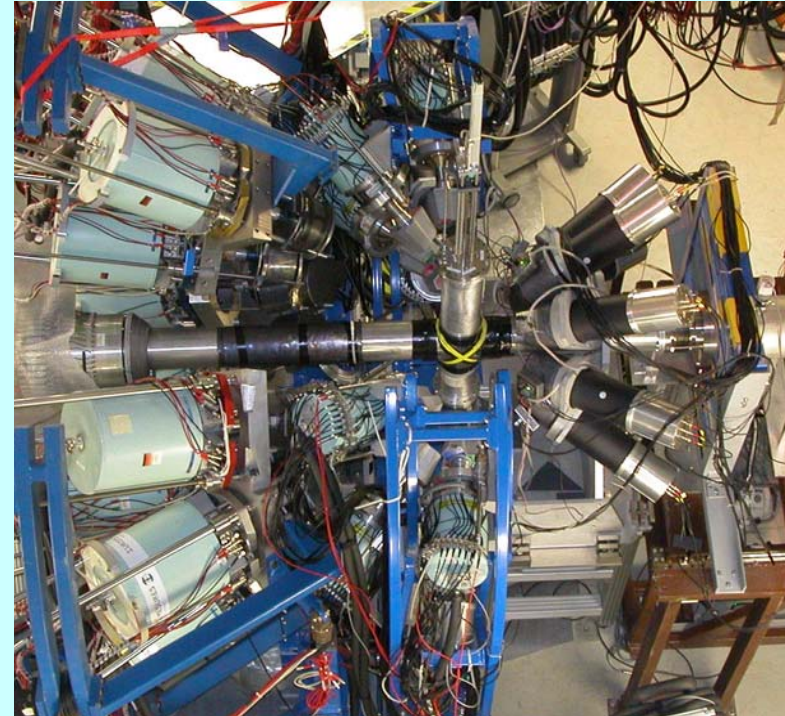
Located at 142° and 88° degrees
Energetic threshold ~ 2 MeV

Miniball **7 HPGe segmented** detectors

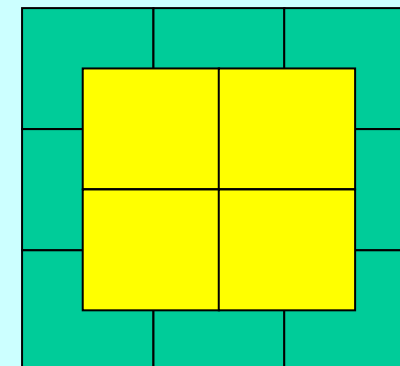
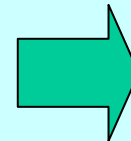
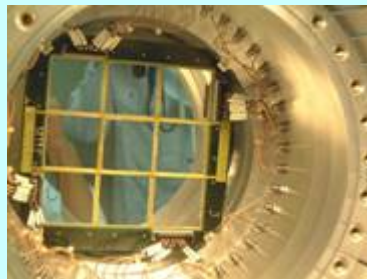
Located at 46° , 60° , 80° , 90° degrees
Energetic threshold ~ 100 keV

Beam identification and **tracking detectors**

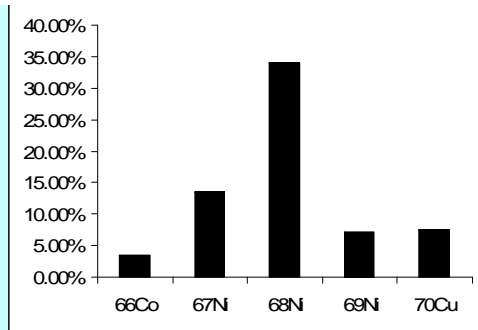
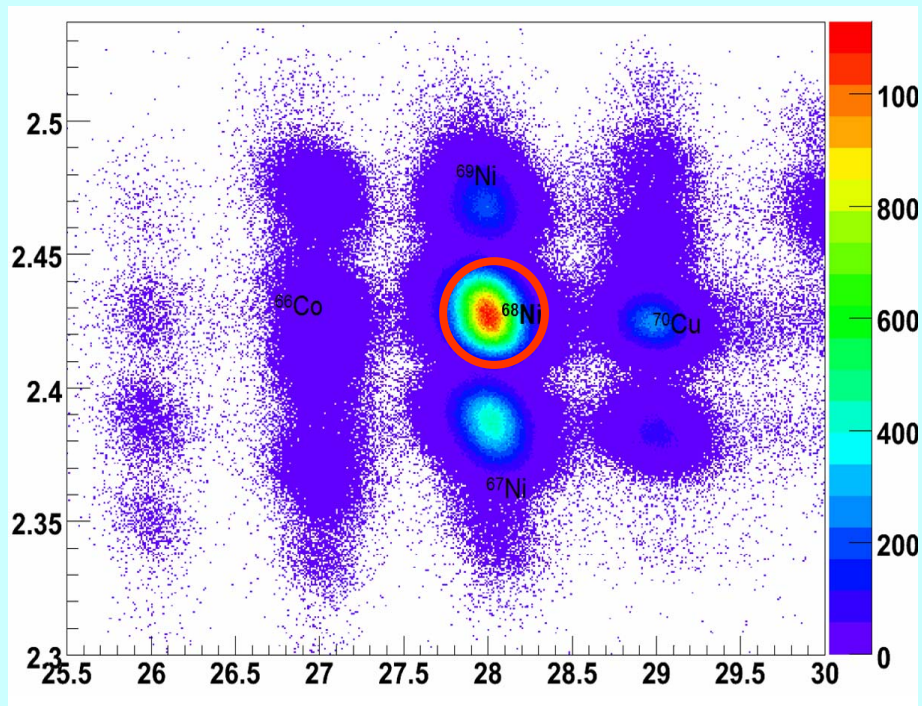
Before and after the target



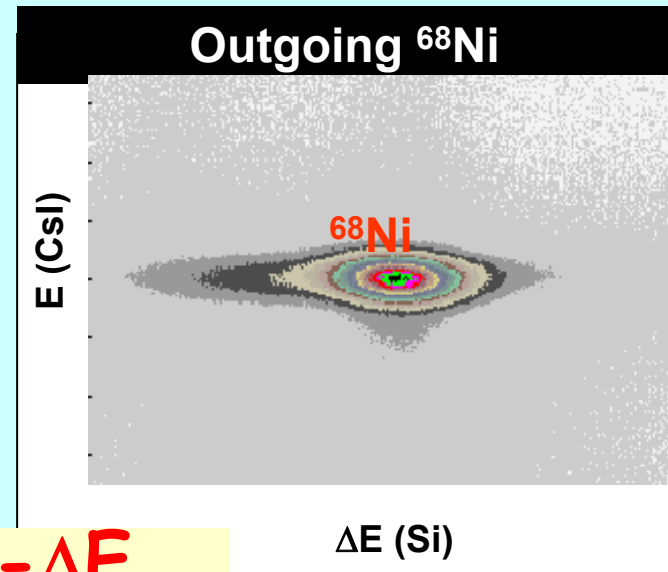
Calorimeter
Telescope
for beam identification
CATE
Position sensitive



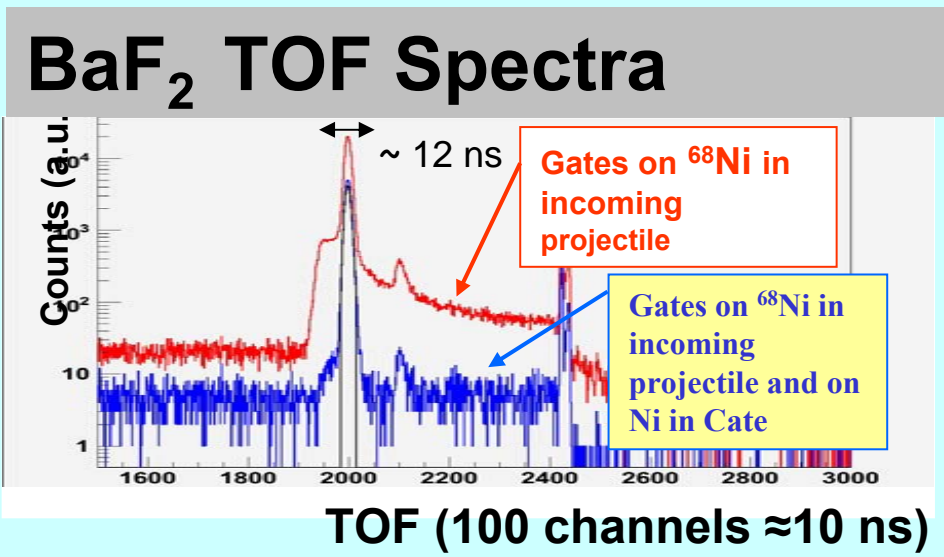
4 CsI
9 Si

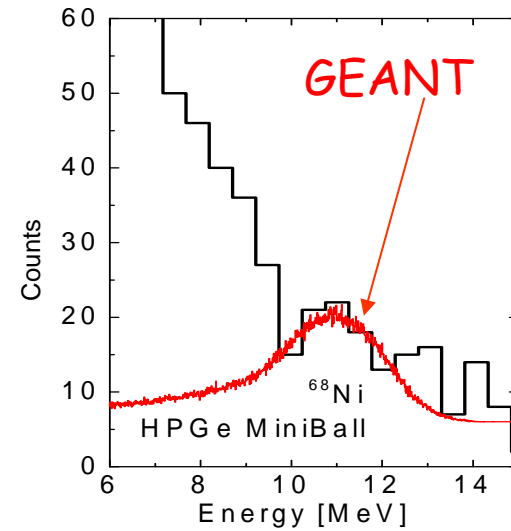
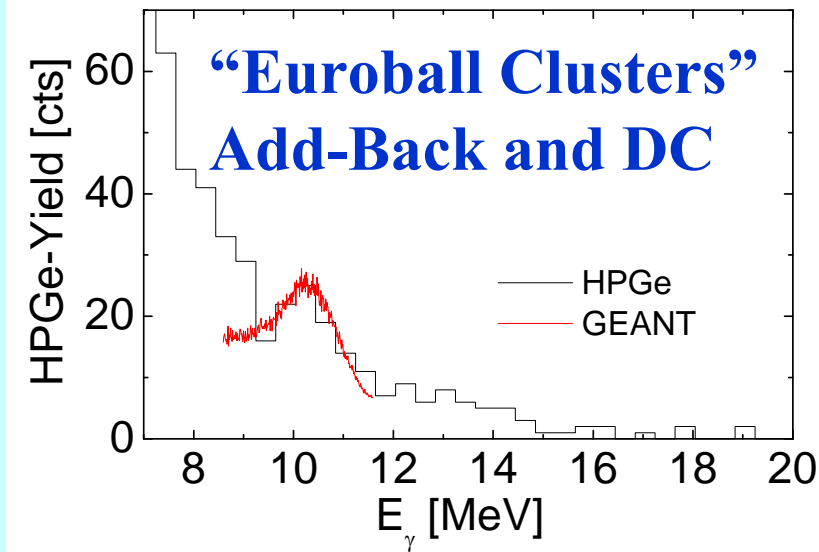


~ 6 Days of effective beam time
 ~ 400 GB of data recorded
 ~ $3 \cdot 10^8$ Events recorded
 ~ $1 \cdot 10^8$ ⁶⁸Ni recorded
 ~ $3 \cdot 10^7$ ' good ⁶⁸Ni events ' recorded



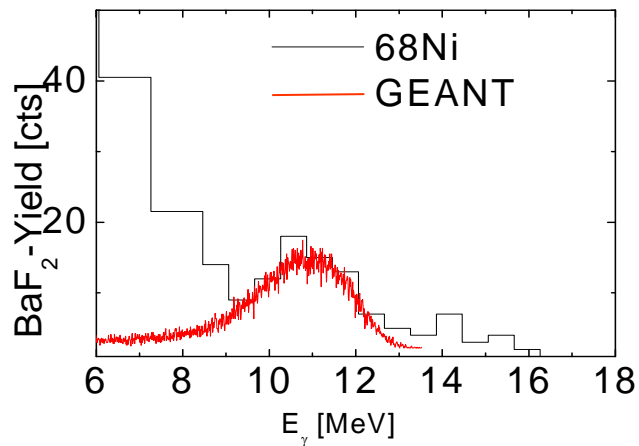
E-ΔE telescopes





Forward : EUROBALL

Center:MINIBALL



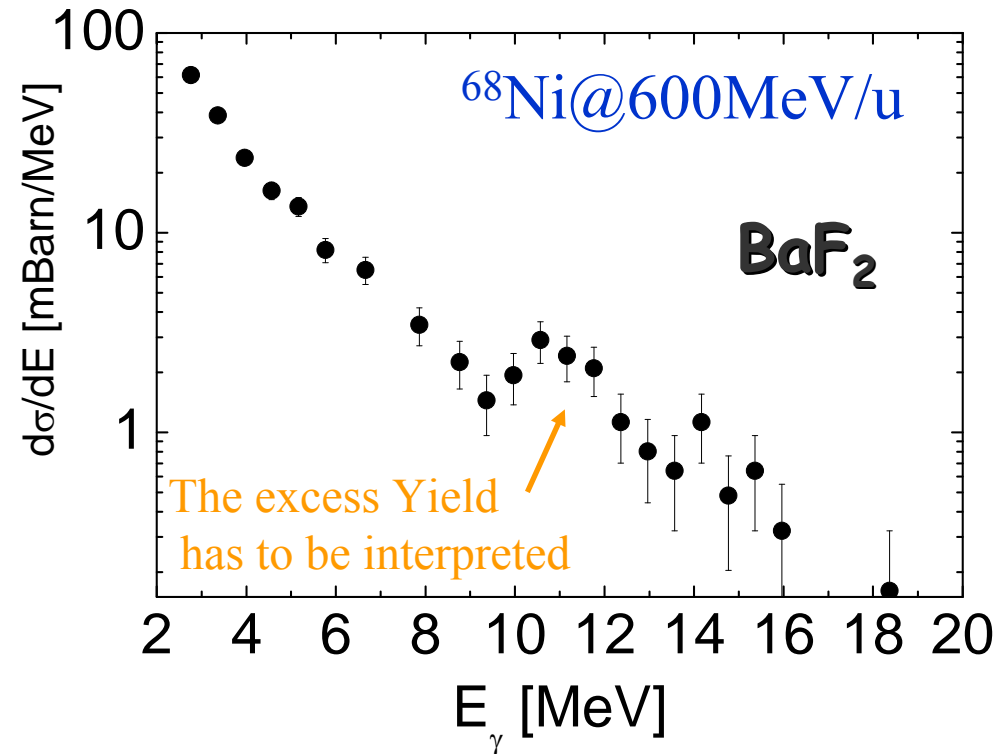
Backward HECTOR

**Structure @ 10.5 MeV
in all detectors**

ANALYSIS and final RESULTS

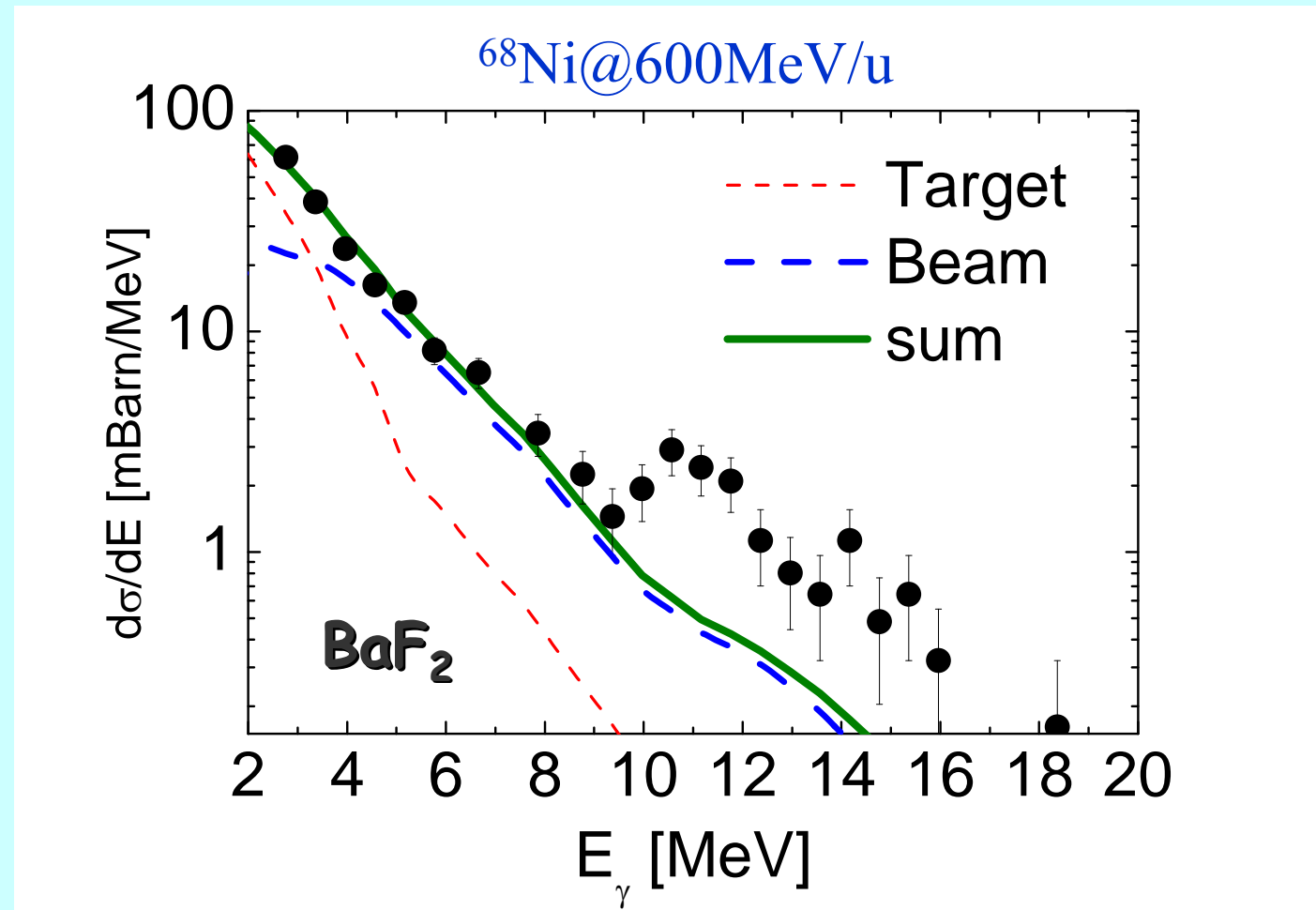
1. Counts into cross Section

$$\sigma_{i \rightarrow f} = \frac{N_{\gamma}}{N_B \cdot N_T \cdot \epsilon}$$



Background ??

2. Evaluation of Background



Statistical model (Cascade) calculation of γ -rays following statistical equilibration of excited **target nuclei** (^{197}Au) and of the excited

beam nuclei (^{68}Ni) folded with RF and in the CM system

[see f.ex. J.Ritman et al. PRL70(1993)533]

3. Data Analysis

Calculate the ground state γ -ray decay
from a GR state following a Coulomb excitation

**! Coulomb excitation probability is directly proportional to
the Photonuclear cross section**

[Eisenberg, Greiner, Bertulani, Alder, Winther, ...]

Coulomb excitation Yield is product of 3 terms:

Virtual photo number, photoabsorption cross section, Branching

$$\frac{d^2\sigma_{C\gamma}}{d\Omega dE_\gamma}(E_\gamma) = \frac{1}{E_\gamma} \frac{dn_\gamma}{d\Omega}(E_\gamma) \sigma_\gamma(E_\gamma) R_\gamma(E_\gamma).$$

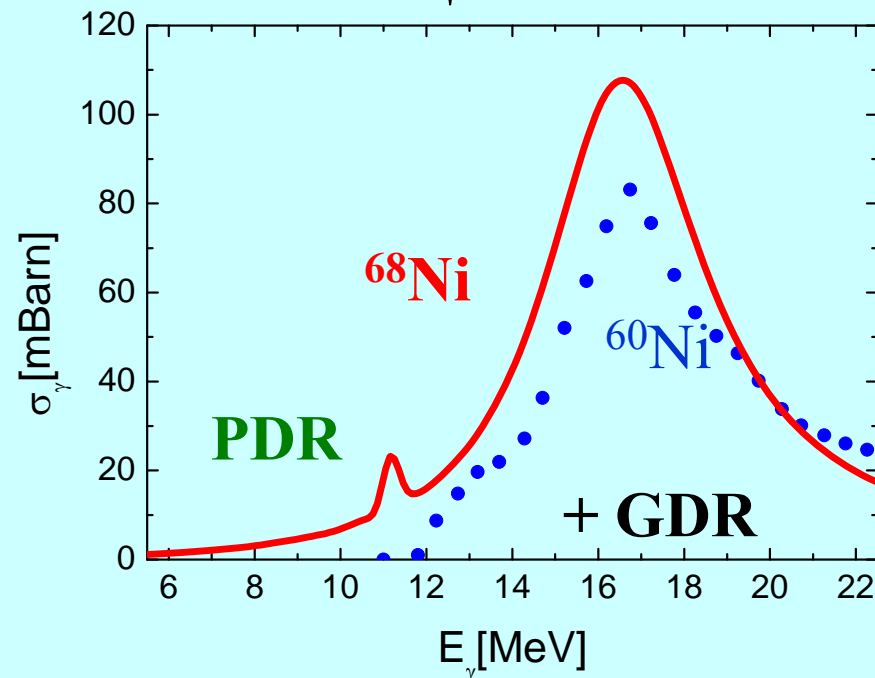
[... Beene, Bortignon,
Bertulani ...]

$$\frac{d^2\sigma_{C\gamma}}{d\Omega dE_\gamma}(E_\gamma) = \frac{1}{E_\gamma} \frac{dn_\gamma}{d\Omega}(E_\gamma) \sigma_\gamma(E_\gamma) R_\gamma(E_\gamma).$$

Photo absorption cross section

the Thomas-Reiche-Kuhn sum rule for $E1$ excitations,

$$\int \sigma_\gamma^{E1}(\epsilon) d\epsilon \simeq 60 \frac{NZ}{A} \text{ MeV mb}$$



$$\frac{d^2\sigma_{C\gamma}}{d\Omega dE_\gamma}(E_\gamma) = \frac{1}{E_\gamma} \frac{dn_\gamma}{d\Omega}(E_\gamma) \sigma_\gamma(E_\gamma) R_\gamma(E_\gamma).$$



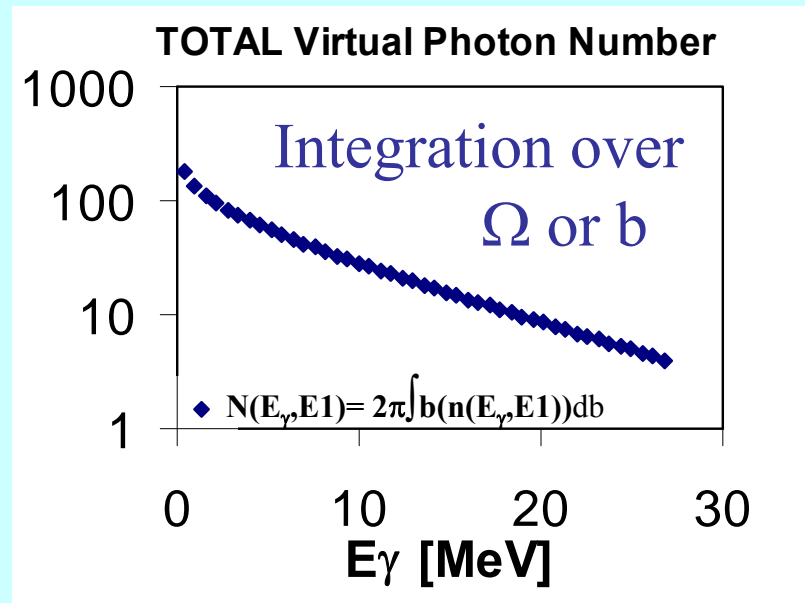
The functions $n_{\pi\lambda}(\varepsilon)$ are called the *virtual photon numbers*, and are given by

$$n_{E1}(b, \varepsilon) = \frac{Z_1^2 \alpha}{\pi^2} \frac{\xi^2}{b^2} \left(\frac{c}{v}\right)^2 \left\{ K_1^2 + \frac{1}{\gamma^2} K_0^2 \right\}$$

= number of equivalent photons
Does NOT depend on the nuclear structure !

Equivalent(virtual)-photon method

Flux of virtual photons per unit area impinging on collision partners.



$$\frac{d^2 \sigma_{C\gamma}}{d\Omega dE_\gamma}(E_\gamma) = \frac{1}{E_\gamma} \frac{dn_\gamma}{d\Omega}(E_\gamma) \sigma_\gamma(E_\gamma) R_\gamma(E_\gamma).$$

Branching Ratio_γ

Two-stage approximation of single sharp states,
 considering the direct GR decay + the compound states :

$$R_\gamma(E_\gamma, \rho_{LD}) = \frac{\Gamma_0^{GR}}{\Gamma^{GR}} + \frac{\Gamma^{GR \downarrow}}{\Gamma^{GR}} C \frac{\langle \Gamma_0^c \rangle}{\langle \Gamma^c \rangle} \quad [\text{Beene, B}^4]$$

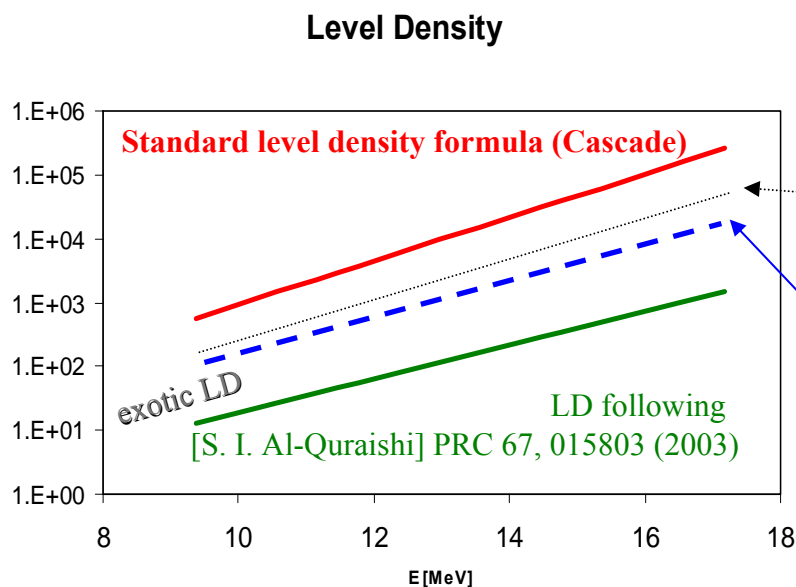
CN

$$\langle \Gamma_0^c \rangle = 2/\pi * \Gamma_{\gamma 0} / \Gamma^{GR} / \rho_{LD}$$

$$\langle \Gamma_{\gamma 0} \rangle = 16\pi/27 * [E/\hbar c]^3 B(E1 \uparrow)$$

$$\langle \Gamma^c \rangle \text{ w. Hauser Feshbach}$$

Level Density

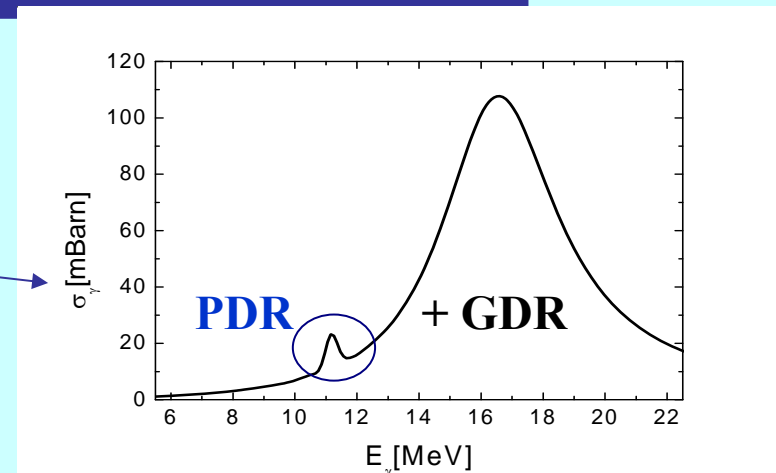


Hartree-Fock-BCS approach (S. Goriely)
 from EMPIRE-Code,
www.nndc.bnl.gov/empire219

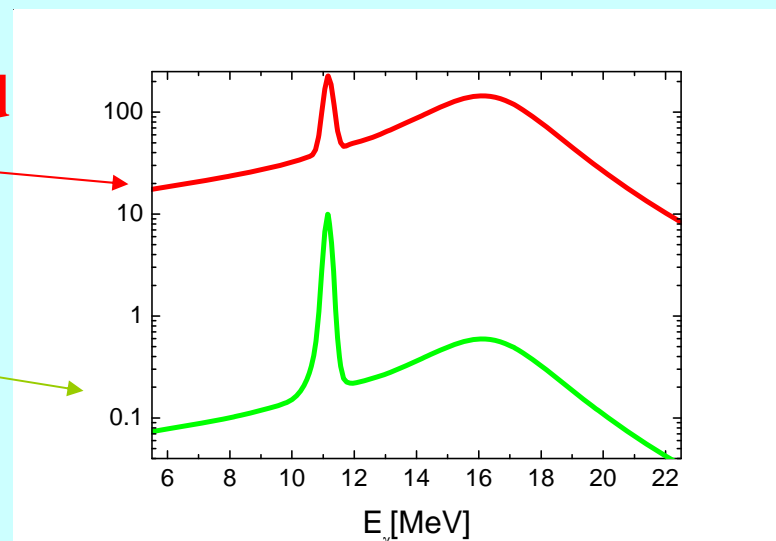
Yoram Alhassid
 [AIP Conf. P 769(2005)1283 and private. com.]
Microscopic nuclear level densities
 (ab initio shell model Monte Carlo method)

$$\frac{d^2\sigma_{C\gamma}}{d\Omega dE_\gamma}(E_\gamma) = \frac{1}{E_\gamma} \frac{dn_\gamma}{d\Omega}(E_\gamma) \sigma_\gamma(E_\gamma) R_\gamma(E_\gamma).$$

photoabsorption cross section



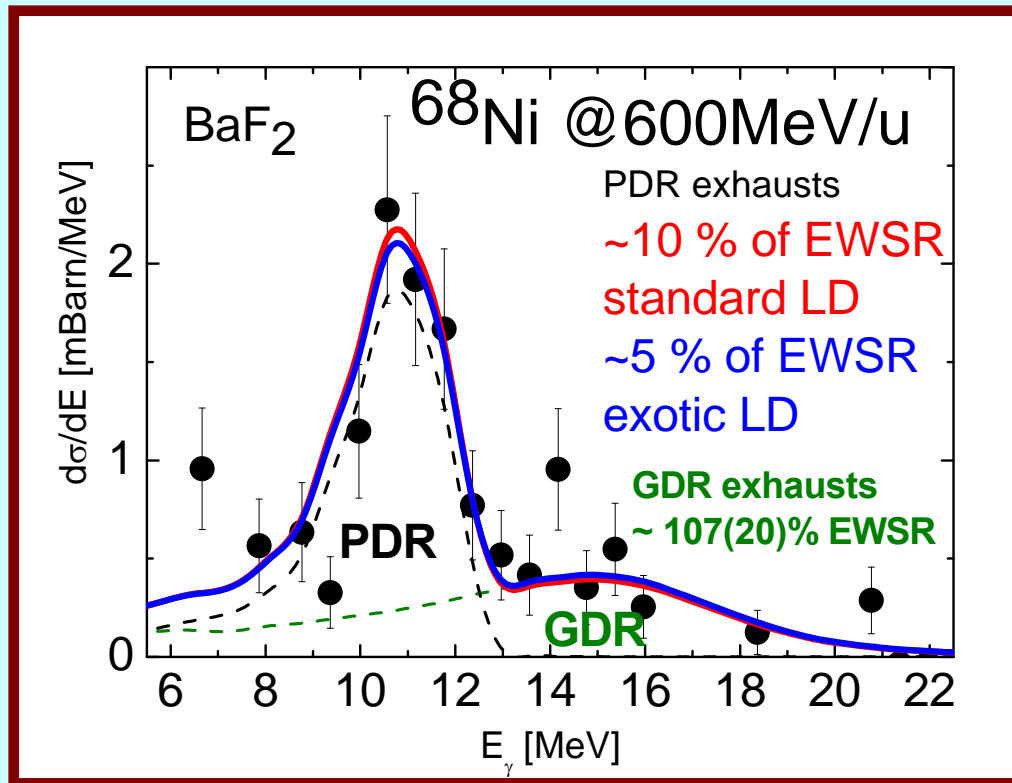
+
Equivalent(virtual)-photon method



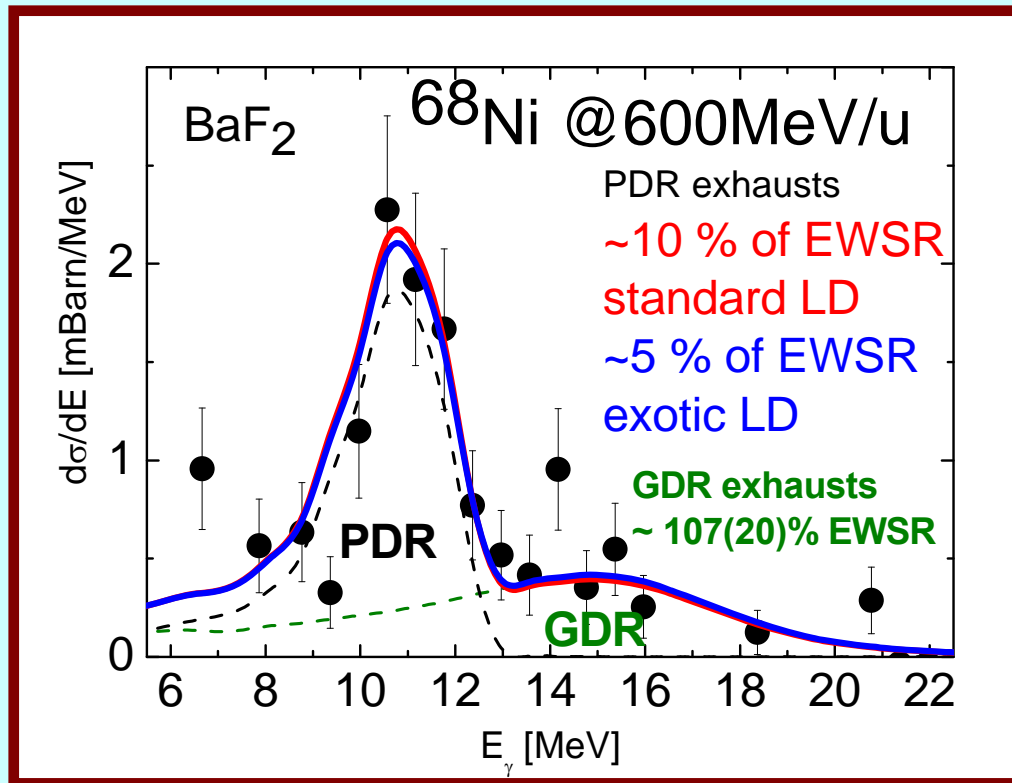
+
Branching

+
Detector Response Function...

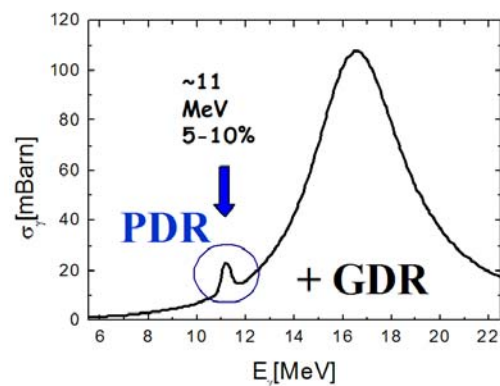
Folded with the **detector response function**
to compare to measured data points ● :



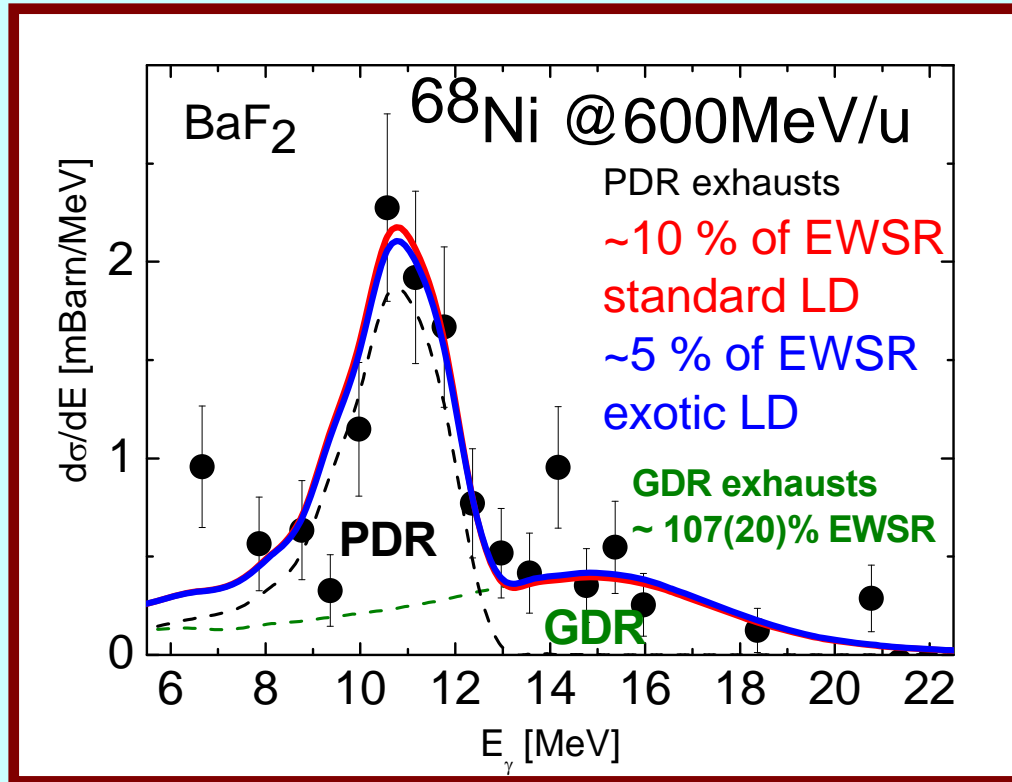
Folded with the **detector response function** to compare to measured data points ● :



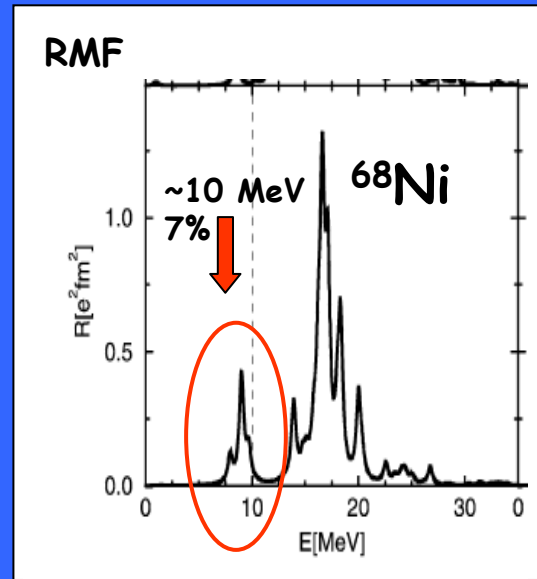
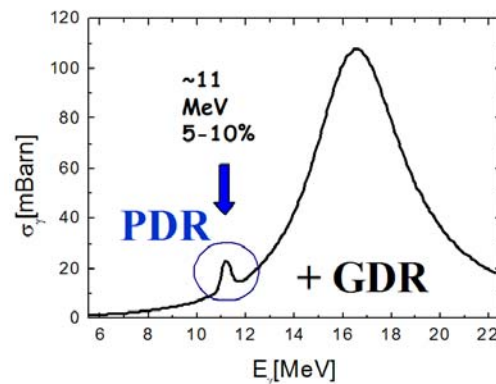
Width of the
PDR is:
 $\Gamma < 500\text{keV}$



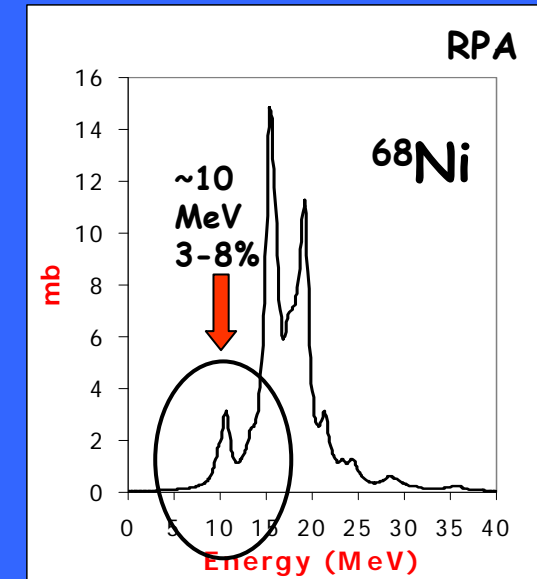
Folded with the **detector response function** to compare to measured data points ● :



Width of the
PDR is:
 $\Gamma < 500\text{keV}$



D. Vretenar et al. NPA 692(2001)496



G. Colo private communications

+ J. Liang et al., PRC75(2007)
frPA: 7-8%:



Conclusions



- Measured high energy γ -rays from Coulex of ^{68}Ni at 600 MeV/u
 - First experiment of this type ever performed (GSI is the only possible Laboratory)
 - We have measured with 3 different detectors a structure around 11 MeV
 - We found an extra strength at 11 MeV with around 5% to 10 % of the EWSR. The error is related by the assumption of the Level density. The theory (RMF and RRPA calculations) predicts 4-8%.
 - The results open new perspectives for other experiments and are very promising for Future measurements especially with high resolution





RISING/PYGMY Collaboration

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